Arcetri Observatory: Astrophysics & Technology





INAF - Arcetri

Dip. di Fisica e Istituto Astronomia Nazionale di Ottica

Villa "II Gioiello" 

- Italian Astronomic Observatories and former CNR Institutes
- 17 Institutes + TNG
- ~ 1200 employees



People

~ 120 employees:
•60 astronomers/technology staff
•30 technical-administrative
•30 PhD/post-doc (15-20 calls/year)
Close cooperation with the Astronomy
group of UniFi



Univ. di Firenze Dip. Di Fisica

Astrophysics: Sun & Plasma Physics

- Historicaly relevant
- Both theoric and instrumental activity
 - Astrophysical plasma
 - Acoustic wave propagation and chromospheric shocks
 - Turbolence
 - Instrumentation (IBIS)



Astrophysics: Solar system & Astrobiology

- Planet formation, origin and diffusion of organics
- Comets, Asteroids
- Planetary explorations
 - Marco Polo-R: sample return mission
 - Stardust: Cometary dust
 - Exomars: Search for life on Mars (life marker chip)





Astrophysics: Star Formation

- Radio, sub-mm e infrared observations
- Chemistry of the interstellar medium
- Protostellar disks-jets







Astrophysics: High Energies

Theoric studies on:
Cosmic rays acceleration
Supernovae remnants
Neutron stars
pulsars



Energies and rates of the cosmic-ray particles



Astrophysics: Galaxies



- Coalescence of cosmologic dark matter halos
- Origin of stars inside dark matter halos
- Diffuse matter (gas, dust, metals)
- Supernovae

Multiwavelength observations

X-ray Optical

Ηα

Infrared

Astrophysics: Active Galactic Nuclei

- Supermassive (10⁷-10⁹ Msun) BH in galactic centers - Quasars
- Co-evolution galaxies-BH
- GR test in strong field regime



Elements formation-evolution

• Matal abundances: cosmology and star formation history



LBT: Large binocular telescope

2 x 8.4m New generation: 1.Adaptive optics 2.Interferometry

4 partners: 1.INAF – I 2.Max-Planck – D 3.Arizona – USA 4.RC – USA

00 808 00



Mount Graham (Arizona)

Arcetri: Adaptive optic systems

- Deformable mirrors (674 actuators, 1kHz)
- Wavefront sensor
- Laser guide star system

Adaptive optics

HST/WFPC3, H band 21min

10"

LBT+FLAO, Ks band 3 min

E-ELT: European extremely-Large Telescope

39m

1.Large collecting area

2.Large angular resolution



ELT Instrumentation

- 1. ELT-IFU (integral-field spectrograph + LTAO+GLAO
- 2. ELT-CAM (high resolution camera) MAORY (adaptiveoptics module)
- 3. ELT-VISIR (mid-IR istrument) +
- 4. ELT-HIRES (high-resolution 400-2500nm spectrograph) +
 SCAO
- 5. ELT-MOS (multi-object spectrograph) + GLAO
- 6. ELT-PCS (spectrograph for exoplanet atmospheres) + XAO



Radioastronomy group: Astrophysics

- Observative: several groups observe and use radio data
 - Galactic: physics and chemistry of molecular clouds; low- and high-mass star formation
 - Extragalactic: atomic and/or molecular gas in nearby galaxies and low-metallicity starbursts
- 10 staff, 2 associate (retired), 2 post doc

Main findings

- IR-dark clouds: detection of bound pre-stellar cores
- Structure/kinematics of jets and disks in low-mass YSOs
- Detection of Keplerian disks around high-mass YSOs
- Chemistry of dense gas: chemical clocks (deuteration, depletion) and pre-biotic molecules (e.g. detection of glycoaldehyde)
- Parallax measurements and Galactic disk structure and kinematics
- Cool dust detection in extreme low-metallicity starbursts
- Detection of weak outflows from low-luminosity AGN

ALMA finds massive, deuterated (prestellar) cores in IRDCs to be bound (Tan + Fontani 2013)



ALMA finds Keplerian disk around highmass (18 M_o) YSO (Sánchez-Monge, Cesaroni, Beltr<u>án, et al. 2013)</u>



study of jet and Keplerian disk in low-mass YSO (Codella et al

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α (J2000)





NGC 1433: ALMA reveals the weakest outflow ever discovered in a lowluminosity AGN (Combes, Garcia-Burillo, Casasola, Hunt et al. 2013). SBS 0335-052: ALMA localizes cool dust in a metal-poor starburst, overthrowing previous estimates of dust-to-gas mass ratios at low metallicity (Hunt et al. 2014).



Radioastronomy group: Technology

Expertise:

- Passive EM optics: feeds, OMT, superconducting filters
- Cryogenics: optical properties of large windows
- Digital signal processing (FPGA-GPU): digital receivers, spectrometers, correlators, pulsar processing
- Spectrum management: Italian liaision for RAS with the National Administration, CRAF Member for INAF

Projects

- Large telescopes: ALMA, SKA
- European Framework programs: Radionet
- Italian telescopes: Medicina, Noto, SRT

Staff

4 research staff, 2 Post-doc

1 technical position, full time, + 4 part time

ASI Technologies for Space Missions





Front-end devices at Q-band



Aluminum Platelet (left) and 3D Laser Synthering (top) OMTs

Filterbank to adapt 2 GHz ALMA IF BW to 62 MHz correlator BW

- Zero cost upgrade of previous correlator design
- 32 fold increase in spectral resolution
- Possibility of mixed resolution modes









ALMA first light spectrum of OriA at 240 GHz

EM analysis and experimental validation of Low Frequency Aperture Arrays











Institute of Electronics, Computer and Telecommunication Engineering











Frequency channelizer for FX correlators

Frequency channelization both in the LFAA beamformer and in the correlators

2¹⁸ channels: Two stage channelization

Filter design and optimization

Effects of quantization



